

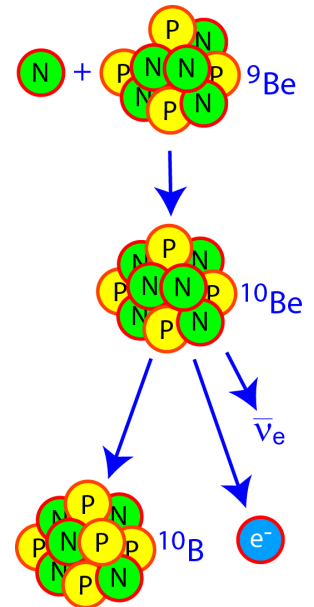
Neutron Capture Processes

This activity uses the “Neutron Capture Process Chart” so students can simulate neutron capture processes and perform their own theoretical calculations!

Estimated time: 20-30 minutes

Many elements are created by fusion in stars, but heavy elements can't be! Those elements may result from neutron capture processes:

1. Free neutrons are created by nuclear reactions in a red giant star.
2. A stable nucleus in the star (Be-9 in the example at right) absorbs a neutron, making a neutron-rich and unstable Be-10.
3. The Be-10 nucleus releases energy/becomes stable by beta-minus decay, turning a neutron into a proton and forming B-10.
4. Thus, a Beryllium nucleus has been turned into a heavier element, Boron!
5. This new stable Boron nucleus might absorb a neutron, and the whole process continues.



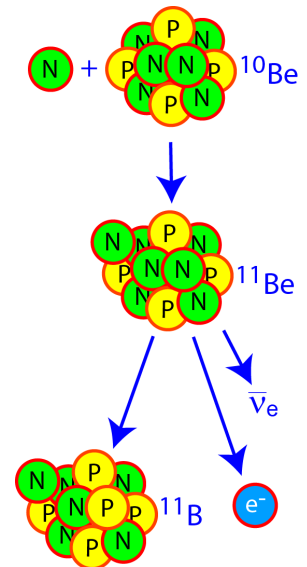
Capture or decay first?

While radioactive, Be-10 is likely to exist at least ten years, long enough to capture a neutron

In a star with loose neutrons around, stable nuclei might absorb them every so often. Unstable nuclei might too, as long as they didn't decay first. Consider the Be-10 above... its half-life is over 1 million years. *If neutrons were abundant enough that a nucleus would normally capture one every ten years, would Be-10 be more likely to capture or decay first?*

Be-11 will almost certainly decay long before the ten-year window is up

A Be-10 nucleus capturing a neutron would become Be-11, which has a half-life of 13.8 seconds. *With the same assumptions above, would it be more likely to capture or decay first?*



The number of neutrons available and the half-life of each isotope determines whether it is more likely to capture a neutron or decay!

Creating a model

Students will perform so many beta-minus decays that they'll forget that beta-plus decay moves in the opposite direction, which can result in very incorrect paths!

Note the “legend” at right: on a chart of the nuclides, neutron capture moves a nucleus to the right, while beta decays go up & left or down & right. **Remember this for the next part!**

Let's construct a simple model of how neutron capture occurs in a red giant star. In our model:

1. Neutrons capture every 10 yr
2. Isotopes with half-lives longer than 10 yr will capture, while isotopes with half-lives shorter than 10 yr will decay.

Using that model, you can make a prediction of the “s-process”, a set of slow neutron capture reactions that may occur in a red giant star!

