

Afterglow Light  
Pattern  
380,000 yrs.

Dark Ages

Development of  
Galaxies, Planets, etc.

Dark Energy  
Accelerated Expansion

# Beyond Standard Model Big-Bang Nucleosynthesis: Fundamental Constants

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Quantum  
Fluctuations

1st Stars  
about 400 million yrs.

Big Bang Expansion

13.7 billion years

- BBN important for synthesis of light elements and evolution of universe
- Element abundances as probes of fundamental physics
- Impact of variations of  $\Lambda_{QCD}$ ,  $G$ ,  $\mu_n$  on BBN

## 1 Background information

- $\Lambda_{QCD}$  and  $m_n$ , deuteron binding energy, neutron lifetime
- $m + n \rightarrow d + \gamma$  reaction rate
- Graviational Constant

## 2 Calculations

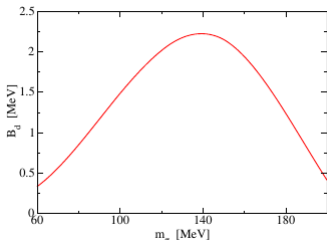
## 3 Results

- Abundance Plots

# Impact of $\Lambda_{QCD}$

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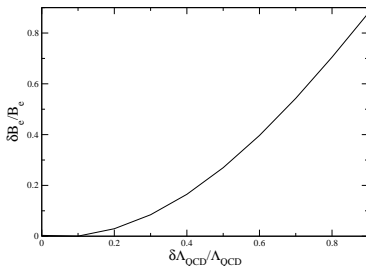
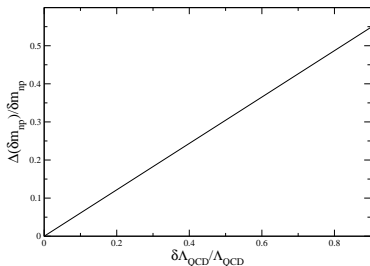
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$B_D$  vs.  $m_\pi$  at NLO  $\chi EFT$ , Savage et al, 2002

Gail McLaughlin et al, 2003

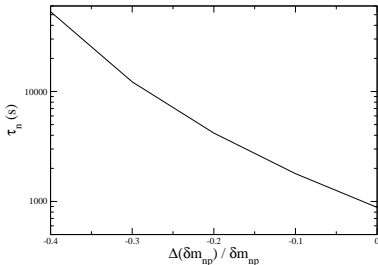
- $\Delta(\delta m_{np}) = -\left(\frac{\Delta\alpha}{\alpha} + \frac{\Delta\Lambda_{QCD}}{\Lambda_{QCD}}\right)\alpha M_{elm}$
- $\frac{\Delta\alpha}{\alpha} \approx 1/30 \frac{\Delta\Lambda_{QCD}}{\Lambda_{QCD}}$
- $\alpha M_{elm} \approx 0.76 \text{ MeV}$
- $\frac{\Delta m_\pi}{m_\pi} = \frac{1}{2} \left(\frac{\Delta\Lambda_{QCD}}{\Lambda_{QCD}}\right)$



# Impact Of $\Lambda_{QCD}$ , $G$

Dependence of neutron lifetime:

$$\tau_n = \frac{(G_F \cos \theta_c)^2}{2\pi^3} m_e^5 (1 + g_A^2) F(\Delta m_{np})$$
$$F(x) = \frac{1}{15} (2x^4 - 9x^2 - 8) \sqrt{x^2 - 1} + x \log(x + \sqrt{x^2 - 1})$$



Friedman Equations:

$$H^2 + \frac{k}{a^2} = \frac{8\pi G}{3} \rho + \frac{\Lambda_{vac}}{3}$$
$$H^2 + \frac{dH}{dt} = -\frac{4\pi G}{3} (\rho + 3p) + \frac{\Lambda_{vac}}{3}$$

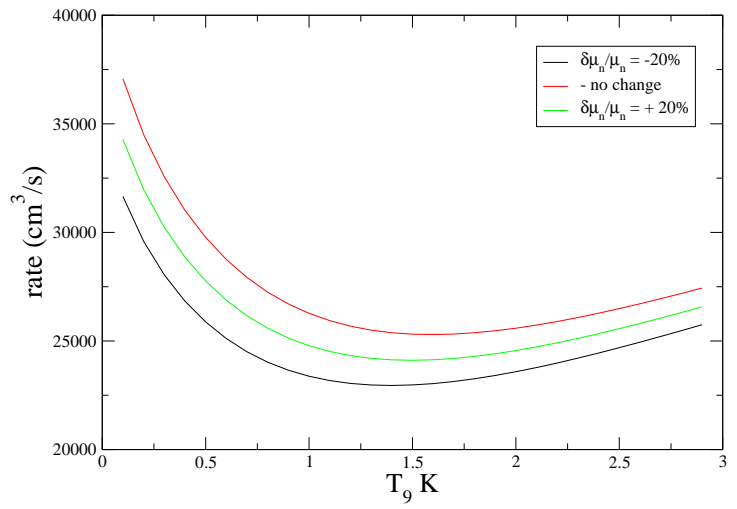
Radiation Dominated,  $k \approx 0$ ,  $\Lambda_{vac} \approx 0$ :

$$\rho = 3p; \quad \rho = \frac{3}{32} \pi G t^2$$
$$\Rightarrow H = \frac{\pi}{2} G t$$
$$\Rightarrow \frac{dH}{dt} = -2H^2$$

# Results : Impact of $\mu_n$

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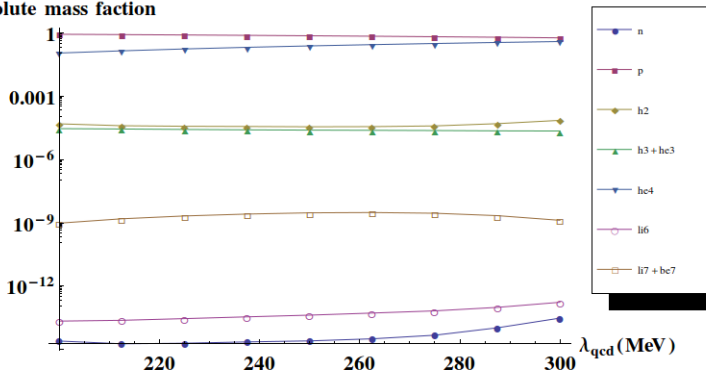


# Results : Mass fraction as function of $\Lambda_{QCD}$ , part 1

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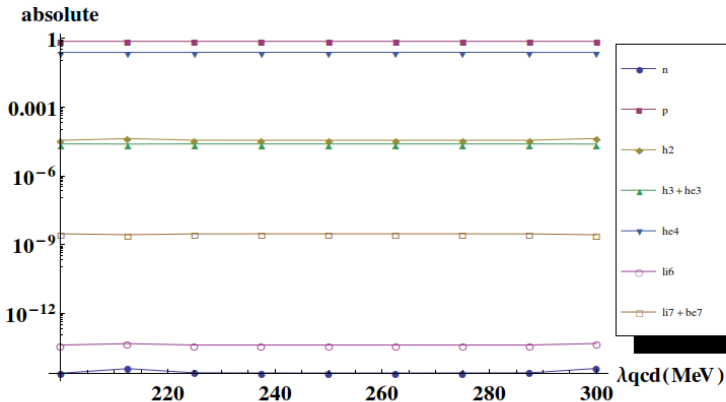
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absolute mass fraction



# Results : Mass fraction as function of $\Lambda_{QCD}$ , part 2

## Neutron Mass Change Only



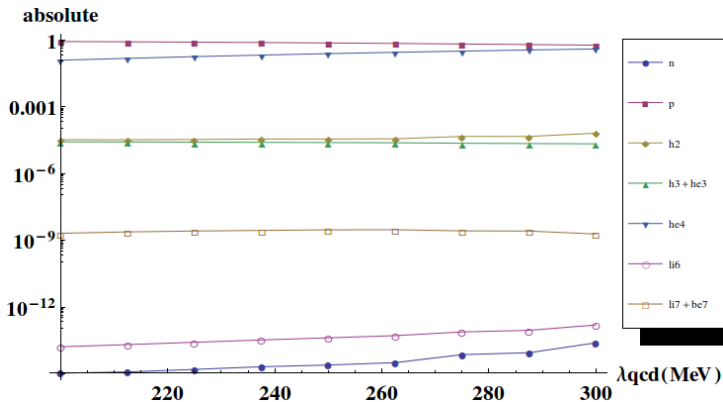


# Results : Mass fraction as function of $\Lambda_{QCD}$ , part 3

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## $\tau_n$ Change only

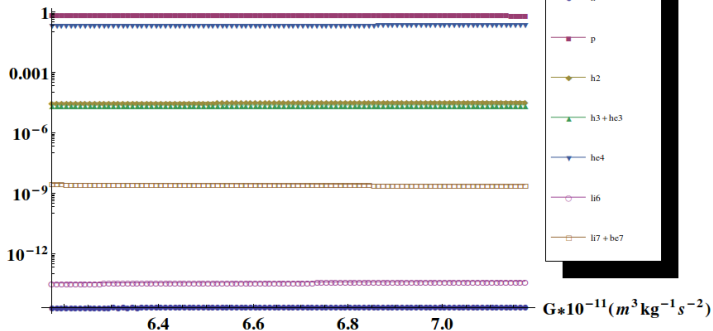


# Results : Mass fraction as function of $G$ , part 1

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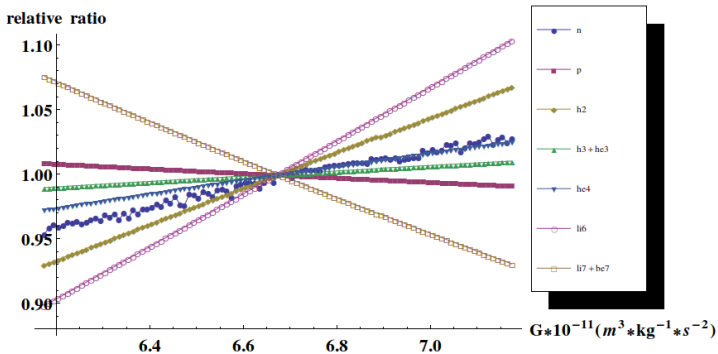
absolute ratio



# Results : Mass fraction as function of $G$ , part 2

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Constants

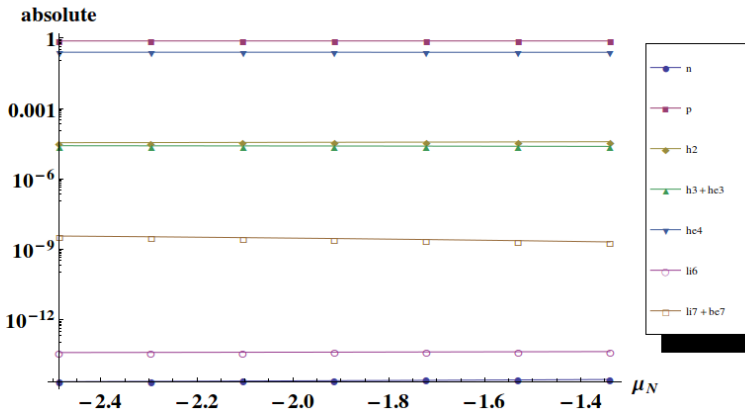
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# Results : Mass fraction as function of $\mu_n$ , part 1

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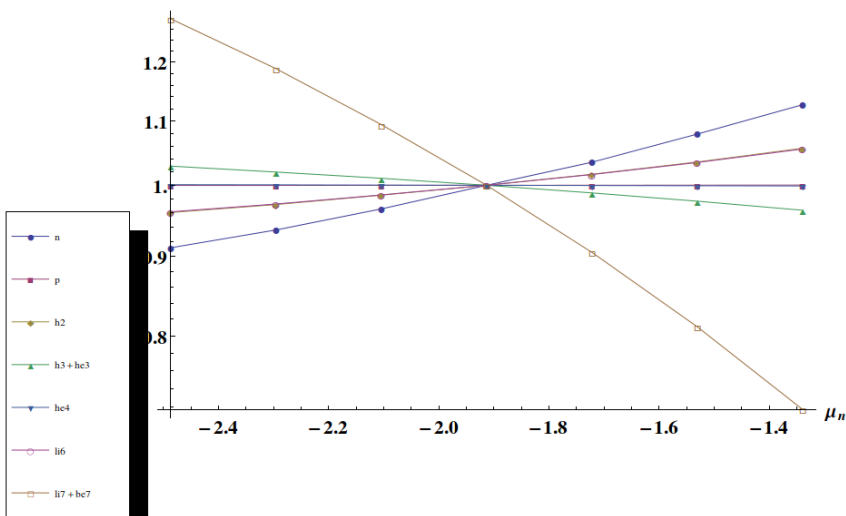


# Results : Mass fraction as function of $\mu_n$ , part 2

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relative mass fraction



- We investigated the dependence of BBN abundances on the values of fundamental constants.
  - $\Lambda_{QCD}$  has a significant impact especially on the Helium-4 abundance, mainly through the change to the neutron decay time.
  - The lithium abundance is highly sensitive to the value of the gravitational constant  $G$ .
  - Changing  $\mu_n$  decreases the rate of production of deuteron
- 
- A change in  $\Lambda_{QCD}$  would probably affect more than the  $p(n, \gamma)d$  rate and free neutron decay. Yet, modifying more interactions is beyond the scope of this 4-day project.